



### **Glacial Cycles**

### Ice Albedo Feedback Model

$$R\frac{\partial T}{\partial t} = \underbrace{Qs\left(y\right)\!\left(1\!-\!\alpha\left(y,\eta\right)\right)\!-\!\left(A\!+\!BT\right)\!+\!C\left(\overline{T}\!-\!T\right)}_{\text{insolation}} \quad \text{albedo} \quad \underbrace{\text{outward}}_{\text{radiation}} \quad \det_{\text{transport}}$$

radiation y = sine of latitude T(y) = annual mean temperature Qs(y) = annual mean insolation

 ${\it Q}={\it global annual mean insolation}$  This equation has a stable equilibrium consisting of polar ice caps.

The latitude of the equilibrium ice boundary and the equilibrium global annual mean temperature are functions of the parameters.

K.K. Tung, Topics in Mathematical Modeling, Princeton (2007), Chapt 8



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$$R\frac{\partial T}{\partial t} = \overline{Qs(y)} \Big( 1 - \alpha(y, \eta) \Big) - \Big( A + BT \Big) + C \Big( \overline{T} - T \Big)$$

#### Idea

Instead of solar forcing (maximum insolation at 65 N latitude), use the global annual mean temperature predicted by the model.

Using Kepler's Laws, we can compute:

$$Q = \frac{Ka^2}{\sqrt{1 - e^2}}$$

$$s(y) = \frac{2}{\pi^2} \int_0^{2\pi} \sqrt{1 - \left(\sqrt{1 - y^2} \sin \beta \cos \gamma - y \cos \beta\right)^2} d\gamma$$

a = semimajor axis e = eccentricity  $\beta = \text{obliquity}$ 



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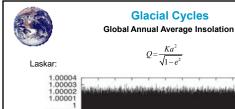
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Note that  $\mathcal{Q}$ , the global annual mean insolation depends only on the semimajor axis and the eccentricity.

Note that s(y), the insolation distribution by latitude, depends only on the obliquity.

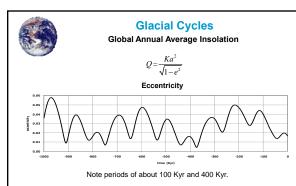
Note that the effect due to precession disappears when averaged over a year.



0.99999 0.99998 0.99997 -250-200-150-100-50 0 50 100 150 200 250 Fig. 11. Variation of the semi-major axis of the Earth–Moon barycenter (in AU) from \_250 to +250 Myr.

Semi major axis does not change much: .005% corresponding to .01% change in global average insolation

J. Laskar, et al (2004) A long-term numerical solution for the insolation quantities of the Earth, Astronomy & Astrophysics 428, 261–285.



As e varies between 0 and 0.06,  $(1-e^2)^{-1/2}$  varies between 1 and 0.0018, or about 0.2%. (Twenty times the effect due to a.)

J. Laskar, et al (2004) A long-term numerical solution for the insolation quantities of the Earth, Astronomy & Astrophysics 428, 261–285.

